# Radiative Equilibrium

## The Condition of Radiative Equilibrium

Modes of Energy Transport

- Radiative
- Convective
- Conduction Not important except in WD's etc
- When all energy is transported by radiation the condition of radiative equilibrium applies.
- Schwarzschild:

 $\tau \land 1$  : Radiative  $\tau > 1$  : Convective

Convection is driven by Hydrogen ionization.

#### **Radiative Equilibrium**

• The Energy Removed from the Beam:  $\int_{0}^{\infty} dv \oint K_{v} I_{v} d\omega = 4\pi \int_{0}^{\infty} K_{v} J_{v} dv$ 

K<sub>v</sub> - total absorption coefficient
Total Energy Replaced In Beam:

$$\oint d\omega \int_0^\infty j_\nu d\nu = 4\pi \int_0^\infty \mathbf{K}_\nu S_\nu d\nu$$

Radiative Equilibrium Demands:

$$\int_0^\infty \mathbf{K}_{\nu} J_{\nu} d\nu = \int_0^\infty \mathbf{K}_{\nu} S_{\nu} d\nu \qquad (*)$$

#### Flux Constancy

 Take the Equation of Transfer and Integrate over all Frequencies and Solid Angles

$$\frac{1}{\rho}\frac{d}{dz}\left(\int_{0}^{\infty} \oint I_{\nu}\mu d\omega\right) = \oint d\omega \int_{0}^{\infty} j_{\nu}d\nu - \int_{0}^{\infty} d\nu \oint K_{\nu}I_{\nu}d\omega$$

But Under Radiative Equilibrium the Right Side = 0!

$$\frac{1}{\rho} \frac{d}{dz} \left( \int_0^\infty \oint I_\nu \mu d\omega \right) = 0$$
$$\frac{dF}{dz} = 0$$

Therefore the Flux is constant with depth!

Total Flux uniquely specified for each atmosphere

#### Consider A Blackbody

A Small Opening in A TE Cavity  $F_{BB}(\nu) = 2\pi \int_{-1}^{1} I_{\nu}(\mu) \mu d\mu$   $= 2\pi B_{\nu}(T) \int_{0}^{1} \mu d\mu$   $= \pi B_{\nu}(T)$ 

Total Flux is the Integral over Frequency:

$$\int_0^\infty \mathcal{F}_{BB}(\nu) d\nu = \pi \int_0^\infty B_{\nu}(T) d\nu = \sigma T^2$$

Define Teff:



#### **One Last Consideration**

Look at (\*) Again

$$\int_{0}^{\infty} K_{\nu} J_{\nu} d\nu = \int_{0}^{\infty} \kappa_{\nu} J_{\nu} d\nu + \int_{0}^{\infty} \sigma_{\nu} J_{\nu} d\nu$$
$$= \int_{0}^{\infty} K_{\nu} S_{\nu} d\nu$$
$$= \int_{0}^{\infty} \kappa_{\nu} B_{\nu} d\nu + \int_{0}^{\infty} \sigma_{\nu} J_{\nu} d\nu$$
$$\therefore$$
$$\int_{0}^{\infty} \kappa_{\nu} J_{\nu} d\nu = \int_{0}^{\infty} \kappa_{\nu} B_{\nu} d\nu$$

Scattering has cancelled out as it puts as much energy into the beam as it takes out!

Radaitive Equilibrium

### HR Diagram

Chapter 13 – "Stellar Rotation" Böhm-Vitense – *Introduction to Stellar Astrophysics* Vol 1

Landolt-Börnstein

